

“On the degree of uncertainty which Local Attraction, if not allowed for, occasions in the Map of a Country, and in the Mean Figure of the Earth as determined by Geodesy; a Method of obtaining the Mean Figure free from ambiguity by a comparison of the Anglo-Gallic, Russian, and Indian Arcs; and Speculations on the Constitution of the Earth’s Crust.” By the Venerable J. H. PRATT, Archdeacon of Calcutta. Communicated by Prof. G. G. STOKES, Sec. R.S. Received October 5, 1863\*.

1. In former communications to the Royal Society I have shown that Local Attraction, owing to the amount it in some places attains, is a more troublesome element to deal with in geodetical operations than had generally been supposed. The Mountains and the Ocean were shown to combine to make the deviation of the plumb-line as much as  $22''\cdot71$ ,  $17''\cdot23$ ,  $21''\cdot05$ ,  $34''\cdot16$  (or quantities not differing materially from them) in the four principal stations of the Great Arc of India between Cape Comorin and the Himmalayas—viz. at Punnee ( $8^{\circ} 9' 31''$ ), Damargida ( $18^{\circ} 3' 15'$ ), Kalianpur ( $24^{\circ} 7' 11''$ ), Kaliana ( $29^{\circ} 30' 48''$ ); and how much these might be increased or lessened by the effect of variations of density in the crust below it was difficult to say. Deviations amounting to at least such quantities as  $7''\cdot61$  and  $7''\cdot87$  were shown to exist in the stations of the Indian Arc, arising from this last cause (see Phil. Trans. 1861, p. 593 (4) and (5)).

M. Otto Struve has lately called attention to similarly important deflections caused by local attraction in Russia—and especially to a remarkable difference of deflection at two stations near Moscow, only about eighteen miles apart, amounting to as much as  $18''$ , which is attributed to an invisible unknown cause in the strata below (see Monthly Notices of the Royal Astronomical Society, April 1862).

2. It is therefore an important inquiry, What degree of uncertainty does Local Attraction, if not allowed for, introduce into the two problems of geodesy, viz. (1) obtaining correct Maps of a country, and (2) determining the Mean Figure of the Earth. I have pointed out the effect on mapping in India, as far as determining the latitudes is concerned, in a former paper. I propose now to consider the subject generally with reference to any country, and taking into account the longitudes as well as the latitudes. The effect upon the determination of the mean figure of the earth I discuss at greater length. By a change, I venture to call it a correction, of Bessel’s method of applying the principle of least squares to the problem, I obtain formulæ for the semiaxes and ellipticity of the Mean Figure involving expressions for the unknown local deflections of the plumb-line at the standard- or reference-stations of the several arcs made use of in the calculation. These formulæ at once show the great degree of uncertainty which an ignorance of the amount of local attraction must introduce into the determi-

\* Read November 26, 1863. See Abstract, vol. xiii. p. 18.

nation of the mean figure. After this I obtain formulæ for the mean figures of the Anglo-Gallic, Russian, and Indian Arcs by the same method, each involving the expression for the unknown local deflection of the plumb-line at the reference-station of the arc concerned. I then show that values of these three unknown deflections can be found which will make the three ellipses which represent the three great arcs almost precisely the same. These deflections are not extravagant quantities, but quite the contrary, being small. I infer, then, that the mean of these three ellipses is in fact the Mean Figure of the Earth, and in this way surmount what was the apparently insuperable difficulty which our ignorance of the amount of local attraction threw in the way of the solution of the problem. The paper concludes with some speculations on the constitution of the earth's crust flowing from the foregoing calculations.

### § 1. *Effect of Local Attraction on Mapping a Country.*

3. In determining differences of latitude and longitude between places by means of the measured lengths which geodesy furnishes, the method of geodesists is to substitute these lengths and the observed middle latitudes in the known trigonometrical formulæ, *using the axes of the MEAN FIGURE of the earth*. It might at first sight appear likely that this would lead to incorrect results, as the actual length measured may lie along a curve different to that of the *mean* form. I propose now to show that no sensible error is introduced by following this course, either in latitude or longitude, if the arc does *not exceed twelve degrees and a half* of latitude, or *fifteen degrees* of longitude in extent.

4. *First. An arc of Latitude.*—Suppose an ellipse drawn in the plane of the meridian through the two stations,  $a$  and  $b$  being its semiaxes;  $c$  the chord joining the stations;  $s$  the length of the arc;  $r$  and  $\theta$ ,  $r'$  and  $\theta'$  polar coordinates to the extremities of the arc from the centre of the ellipse;  $l$  and  $l'$  their observed latitudes;  $\lambda$  the amplitude of the arc;  $m$  its middle latitude: then we have the following formulæ, neglecting the square of the ellipticity ( $\epsilon$ ),

$$\begin{aligned} s &= \frac{1}{2}(a+b)\lambda - \frac{2}{2}(a-b) \sin \lambda \cos 2m, \\ r &= a(1 - \epsilon \sin^2 l), \quad r' = a(1 - \epsilon \sin^2 l'), \quad \tan \theta = (1 - 2\epsilon) \tan l, \\ \tan \theta' &= (1 - 2\epsilon) \tan l'. \end{aligned}$$

Now

$$\begin{aligned} c^2 &= r^2 + r'^2 - 2rr' \cos(\theta - \theta') = 2rr' \{1 - \cos(\theta - \theta')\} + (r - r')^2 \\ &= 2rr' \{1 - \cos(\theta - \theta')\}. \end{aligned}$$

By expanding the formulæ for  $\tan \theta$  and  $\tan \theta'$ , we have

$$\theta = l - \epsilon \sin 2l, \quad \theta' = l' - \epsilon \sin 2l',$$

$$\begin{aligned} \therefore \theta - \theta' &= l - l' - \epsilon(\sin 2l - \sin 2l') = l - l' - 2\epsilon \sin(l - l') \cos(l + l') \\ &= \lambda - 2\epsilon \sin \lambda \cos 2m; \end{aligned}$$

$$\begin{aligned} \therefore 1 - \cos(\theta - \theta') &= 1 - \cos \lambda - 2\epsilon \sin^2 \lambda \cos 2m \\ &= 2 \sin^2 \frac{\lambda}{2} \{1 - 2\epsilon(1 + \cos \lambda) \cos 2m\}. \end{aligned}$$

Also

$$\begin{aligned}
 rr' &= a^2 \{1 - \epsilon(\sin^2 l + \sin^2 l')\} = a^2 \left\{1 - \frac{\epsilon}{2}(2 - \cos 2l - \cos 2l')\right\} \\
 &= a^2 \{1 - \epsilon(1 - \cos \lambda \cos 2m)\}; \\
 \therefore c^2 &= 4a^2 \sin^2 \frac{\lambda}{2} \left\{1 - \epsilon \{1 + (2 + \cos \lambda) \cos 2m\}\right\}; \\
 \therefore \sin \frac{\lambda}{2} &= \frac{c}{2a} \left\{1 + \frac{\epsilon}{2} \{1 + (2 + \cos \lambda) \cos 2m\}\right\}; \\
 \therefore \frac{\lambda}{2} &= \sin^{-1} \frac{c}{2a} + \frac{\epsilon}{2} \left\{1 + (2 + \cos \lambda) \cos 2m\right\} \frac{c}{\sqrt{4a^2 - c^2}} \\
 &= \sin^{-1} \frac{c}{2a} + \frac{\epsilon}{2} \left\{1 + (2 + \cos \lambda) \cos 2m\right\} \tan \frac{\lambda}{2}.
 \end{aligned}$$

Hence by the first formula,

$$\begin{aligned}
 s &= a \left(1 - \frac{\epsilon}{2}\right) \lambda - \frac{3}{2} a \epsilon \sin \lambda \cos 2m \\
 &= a(2 - \epsilon) \sin^{-1} \frac{c}{2a} + a \epsilon \{1 + (2 + \cos \lambda) \cos 2m\} \tan \frac{\lambda}{2} - \frac{3}{2} a \epsilon \sin \lambda \cos 2m \\
 &= (a + b) \sin^{-1} \frac{c}{2a} + (a - b) \left\{1 + \frac{1}{2}(1 - \cos \lambda) \cos 2m\right\} \tan \frac{\lambda}{2}.
 \end{aligned}$$

Taking the variation of  $s$  with respect to  $a$  and  $b$ , considering  $c$  as constant, and  $\lambda$  and  $m$  also constant, occurring as they do only in small terms, we shall have the difference in length of two arcs joining the stations and belonging to different ellipses, only having their axes parallel. Hence

$$\begin{aligned}
 \delta s &= (\delta a + \delta b) \sin^{-1} \frac{c}{2a} - \frac{a+b}{a} \frac{c \delta a}{\sqrt{4a^2 - c^2}} \\
 &\quad + (\delta a - \delta b) \left\{1 + \frac{1}{2}(1 - \cos \lambda) \cos 2m\right\} \tan \frac{\lambda}{2}.
 \end{aligned}$$

Since the terms are small, we may use the first approximate value for  $c$  and  $b$ ;

$$\begin{aligned}
 \therefore \delta s &= (\delta a + \delta b) \frac{\lambda}{2} - 2 \tan \frac{\lambda}{2} \delta a + (\delta a - b) \left\{1 + \frac{1}{2}(1 - \cos \lambda) \cos 2m\right\} \tan \frac{\lambda}{2} \\
 &= (\delta a + \delta b) \left(\frac{\lambda}{2} - \tan \frac{\lambda}{2}\right) + (\delta a - \delta b) \frac{1}{2} \tan \frac{\lambda}{2} (1 - \cos \lambda) \cos 2m \\
 &= (\delta a + \delta b) P + (\delta a - \delta b) Q \cos 2m,
 \end{aligned}$$

where

$$\begin{aligned}
 P &= \frac{1}{2} \lambda - \tan \frac{1}{2} \lambda, \text{ and } Q = \frac{1}{2} \tan \frac{1}{2} \lambda (1 - \cos \lambda) \\
 &= (P + Q \cos 2m) \delta a + (P - Q \cos 2m) \delta b.
 \end{aligned}$$

I will find the values of  $\delta a$  and  $\delta b$  which will satisfy this equation and make  $\delta a^2 + \delta b^2$  a minimum.

$$\delta a^2 + \left( \frac{\delta s - (P + Q \cos 2m) \delta a}{P - Q \cos 2m} \right)^2 = \text{a minimum};$$

$$\therefore \{ (P - Q \cos 2m)^2 + (P + Q \cos 2m)^2 \} \delta a = (P + Q \cos 2m) \delta s;$$

$$\therefore \delta a = \frac{P + Q \cos 2m}{P^2 + Q^2 \cos^2 2m} \frac{\delta s}{2}, \quad \delta b = \frac{P - Q \cos 2m}{P^2 + Q^2 \cos^2 2m} \frac{\delta s}{2},$$

$$\delta a^2 + \delta b^2 = \frac{1}{P^2 + Q^2 \cos^2 2m} \frac{\delta s^2}{2}.$$

This is least when  $m=0$  and  $90^\circ$ ; then

$$\delta a = \frac{P+Q}{P^2+Q^2} \frac{\delta s}{2}, \quad \delta b = \frac{P-Q}{P^2+Q^2} \frac{\delta s}{2}, \quad \delta a \sim \delta b = \frac{Q \delta s}{P^2+Q^2}.$$

Let one of the two ellipses be equal to the mean ellipse of the earth's figure,  $a$  and  $b$  being the semiaxes, and  $\delta a$  and  $\delta b$  the excess (or defect, if negative) of the semiaxes of the other ellipse. The first ellipse is not necessarily the mean ellipse itself, but is only equal to it in dimensions, and parallel to it in position; for the actual arc may lie above or below the mean ellipse. The result of this is, that the arc of the mean ellipse which corresponds with  $s$  of the actual arc will not necessarily have precisely the same middle latitude, although the chord  $c$  is of the same length. But as the middle latitude will differ only by a quantity of the order of the ellipticity, this difference will not appear in the result, because we neglect the square of the ellipticity.

I will now make the extravagant supposition that the ellipse to which the arc actually belongs deviates from the form of the mean ellipse so much that  $\delta a \sim \delta b = 13$  miles, the whole compression of the earth's figure. On this supposition I will find how large the arc may be so as not to produce a difference in length greater than  $1''$ .

Put  $\delta a \sim \delta b = 13$ ,  $\delta s = 1'' = 0.0193$  mile ( $1^\circ$  being  $69.5$  miles),

$$\therefore (P^2 + Q^2) \div Q = 0.0193 \div 13 = 0.0015,$$

or

$$\left( \frac{\lambda}{2} - \tan \frac{\lambda}{2} \right)^2 + \frac{1}{4} \tan^2 \frac{\lambda}{2} (1 - \cos \lambda)^2 = 0.00075 \tan^2 \frac{\lambda}{2} (1 - \cos \lambda).$$

A slight inspection of this equation shows that  $\lambda$  must be small. Expand in powers of  $\lambda$ ; then

$$\left( \frac{1}{9} + 1 \right) \left( \frac{\lambda}{2} \right)^3 = 0.0015, \text{ or } \left( \frac{\lambda}{2} \right)^3 = 0.00135;$$

$$\therefore \lambda = 0.22 \text{ (in arc)} = 0.22 \times 57.3 \text{ (in degrees)} = 12.6.$$

This shows that in an arc of meridian as much as twelve degrees and a half in length, it would require a departure from the mean ellipse equal to the whole actual compression of the pole of the earth in order to produce so slight a difference in the length as  $1''$ . Hence we may conclude that the difference in length between the mean arc and the actual arc, joining any two places on the same meridian, is an insensible quantity, since an extravagant hypothesis regarding the departure of the form from the mean form will not produce a difference in length of more than  $1''$ . This being the

case, the differences of latitudes calculated from the measured arcs of meridian with the mean axes, as is done in the Survey operations, will come out free from any effects which local attraction can produce, as that attraction can never be capable of causing so great a distortion in the measured arcs as I have supposed for the sake of calculation. The absolute latitude, however, of the station which fixes the arc on the map will be unknown to the extent of the deviation of the plumb-line caused by local attraction at that place.

5. *Second. An Arc of Longitude.*—Let  $S$  be the length of the arc,  $l$  the latitude,  $L$  the longitudinal amplitude or the difference of the longitudes of its extremities,  $c$  the chord. Then

$$S = L \cos l \{a + (a-b) \sin^2 l\}, \quad c = 2 \cos l \{a + (a-b) \sin^2 l\} \sin \frac{1}{2} L.$$

When  $a$  and  $b$  vary,  $c$  and  $l$  remain constant, but  $S$  and  $L$  vary. Hence

$$\begin{aligned} \delta S &= \delta L \cos l \{a + (a-b) \sin^2 l\} + L \cos l \{\delta a + (\delta a - \delta b) \sin^2 l\} \\ 0 &= \{a + (a-b) \sin^2 l\} \cos \frac{1}{2} L \delta L + 2\{\delta a + (\delta a - \delta b) \sin^2 l\} \sin \frac{1}{2} L. \end{aligned}$$

By eliminating  $\delta L$  from these,

$$\delta S = \left( L - 2 \tan \frac{1}{2} L \right) \cos l \{\delta a + (\delta a - \delta b) \sin^2 l\};$$

$$\therefore \delta a + (\delta a - \delta b) \sin^2 l = \frac{\delta S}{(L - 2 \tan \frac{1}{2} L) \cos l} = n, \text{ suppose.}$$

I will, as before, find the values of  $\delta a$  and  $\delta b$  which satisfy this equation, and make  $\delta a^2 + \delta b^2$  a minimum.

$$\sin^4 l \delta a^2 + \{(1 + \sin^2 l) \delta a - n\}^2 = \text{a minimum};$$

$$\therefore \{\sin^4 l + (1 + \sin^2 l)^2\} \delta a = n(1 + \sin^2 l);$$

$$\therefore \delta a = \frac{(1 + \sin^2 l)n}{\sin^4 l + (1 + \sin^2 l)^2}, \quad \delta b = \frac{-\sin^2 l \cdot n}{\sin^4 l + (1 + \sin^2 l)^2};$$

$$\therefore \delta a^2 + \delta b^2 = \frac{n^2}{\sin^4 l + (1 + \sin^2 l)^2} = \frac{\delta S^2}{\cos^2 l \{\sin^4 l + (1 + \sin^2 l)^2\} \{L - 2 \tan \frac{1}{2} L\}^2}$$

This is least when  $\cos^2 l \{\sin^4 l + (1 + \sin^2 l)^2\}$  is greatest, or when  $l=0$ ; then

$$\delta a = n, \quad \delta b = 0, \quad \delta a - \delta b = n = \frac{\delta S}{L - 2 \tan \frac{1}{2} L}.$$

Now put  $\delta a \sim \delta b = 13$  miles,  $\delta S = \text{arc } 1''$  of a great circle  $= 0.0193$  mile;

$$\therefore L - 2 \tan \frac{1}{2} L = 0.0193 \div 13 = 0.0015.$$

This shows that  $L$  must be small: expanding, we have

$$L^3 = 0.018, \quad L = 0.262 \text{ (in arc)} = 0.262 \times 57.3 \text{ (in degrees)} = 15^\circ.$$

We can reason from this, as before, that the differences of longitudes will be accurately found by using the measured arcs of longitude and the mean axes, if the arcs are not longer than  $15^\circ$ . Now arcs of this length, and of

the length determined in paragraph 4 for latitudes, are never used in survey operations: the great arcs are always divided into much smaller portions. Hence the maps constructed from geodetic operations will always be relatively correct in themselves; but the precise position of the map on the terrestrial spheroid will be unknown by the amount of the unknown deflection of the plumb-line in latitude and longitude at the place which fixes the map. In India the effect of the Himalaya Mountains and the Ocean, taken alone, would throw out the map by nearly half a mile. The calculations, however, which I give in the next two sections of this paper, show that the effect of variations in the density of the crust below almost entirely counteracts that of the mountains and ocean at Damargida in latitude  $18^{\circ} 3' 15''$ , and the displacement of the map is almost insensible if fixed by that station. If fixed by the observed latitude of any other station, the map will be out of its place by the local deflection of the plumb-line at that station. This, in the Indian Great Arc, does not exceed one-thirteenth of a mile at any of the stations where the latitude has been observed. It appears also from those calculations, that, except in places evidently situated in most disadvantageous positions, the local attraction is rarely of any considerable amount.

§ 2. *Effect of Local Attraction on the Determination of the Mean Figure of the Earth.*

6. The mean radius of the earth is nearly 20890000 feet, the ellipticity is nearly  $\frac{1}{300}$ , and it is found convenient to put the semiaxes of the earth's figure under the form

$$\left. \begin{aligned} \frac{a+b}{2} &= \left(1 - \frac{u}{10000}\right) 20890000 = 20890000 - 2089 u \text{ feet,} \\ \frac{a-b}{2} &= \frac{1}{600} \left(1 - \frac{u}{10000} + \frac{v}{50}\right) 20890000 = \frac{1}{600} \left(\frac{a+b}{2} + 417800 v\right); \end{aligned} \right\} (1)$$

$u$  and  $v$  are quantities to be determined, and the squares and product of these may be neglected.

Also, ellipticity  $= \frac{a-b}{a} = \frac{1}{300} \left(1 + \frac{v}{50}\right)$ .

The arcs which are actually measured in geodesy do not necessarily belong to precisely the same ellipse: in fact those arcs may not precisely belong to any ellipse. Suppose one of these measured arcs is laid along the ellipse of which the axes are given above, and that, small corrections  $x$  and  $x'$  being added to the observed latitudes of its extremities, the arc with its corrected latitudes exactly fits this ellipse. Then  $x' - x$  may be expressed in the form  $m + \alpha u + \beta v$ , where  $m$ ,  $\alpha$ , and  $\beta$  are functions of the measured length, the observed latitudes, and numerical quantities. Let this be done for all the arcs which have been measured and their subdivisions. I shall take the eight arcs used in the chapter on the Figure of the Earth in the Volume of the British Ordnance Survey; viz. the Anglo-Gallic,

Russian, Indian II. (or Great Arc), Indian I., Prussian, Peruvian, Hanoverian, and Danish Arcs. Suppose

$$m_1 + \alpha_1 u + \beta_1 v + \dots x_1, \quad m'_1 + \alpha'_1 u + \beta'_1 v + x_1,$$

are the corrections of the latitudes of the extremities of the subdivisions of the Anglo-Galic Arc,  $x_i$  being the correction for the standard or reference station in this Arc. Similarly, let

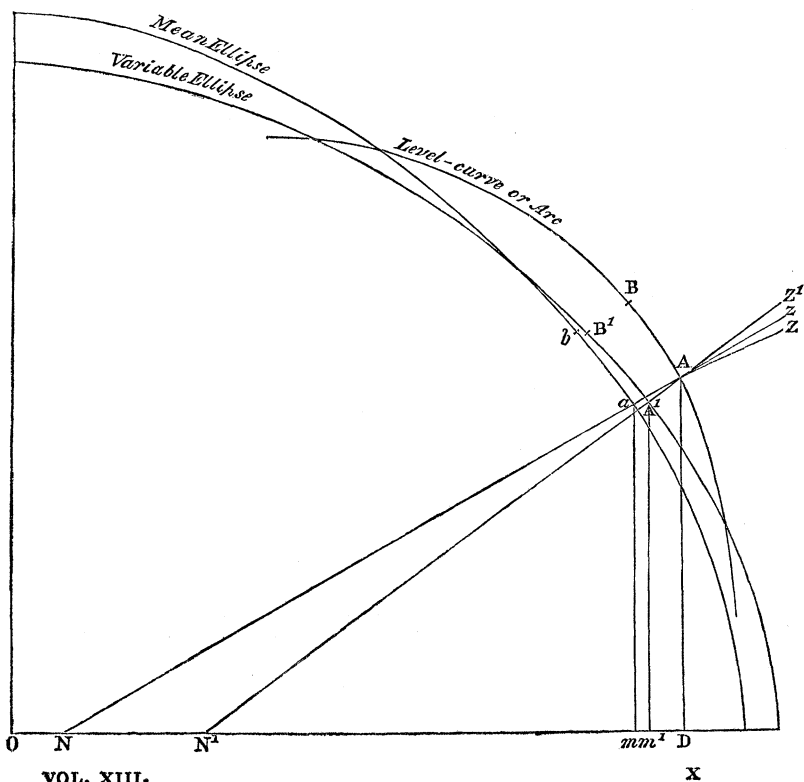
$$m_2 + \alpha_2 u + \beta_2 v + x_2, \quad m'_2 + \alpha'_2 u + \beta'_2 v + x_2, \dots$$

$$m_3 + \alpha_3 u + \beta_3 v + x_3, \quad m'_3 + \alpha'_3 u + \beta'_3 v + x_3, \dots$$

be the corrections for the divisions of the other Arcs.

Then the values of  $u$  and  $v$  which give the most likely form are those which make the sum of the squares of all these corrections a minimum. The sum of the squares will involve  $u$  and  $v$ , and also eight quantities  $x_1 \dots x_8$ . The usual course is to regard, not only  $u$  and  $v$ , but  $x_1 \dots x_8$  as independent variables, and to differentiate the sum of the squares with regard to each of them in succession, and so obtain as many equations as quantities to be determined.

7. This mode of proceeding is, I conceive, erroneous; as I shall now endeavour to show. The corrections  $x_1, \dots, x_8$  are not properly indepen-



dent variables, but are functions of  $u$  and  $v$ , and of the deflections produced by local attraction. In the preceding diagram the plane of the paper is the plane of the meridian in which the arc, of which  $AB$  is one section, has been geodetically measured.  $A$  is the reference-station of the several portions of the whole arc.  $AZ$  is the vertical at  $A$  in which the plumb-line hangs. The two curves, of which  $A'B'$  and  $ab$  are portions, are a variable ellipse and the mean ellipse having the same centre  $O$  and their axes in the same lines, the mean ellipse being what the variable ellipse becomes when the values are substituted for  $u$  and  $v$  which make the sum of the squares of the errors a minimum:  $Z'A A'N'$  and  $zA aN$  are normals through  $A$  to these two ellipses;  $AD$ ,  $A'm'$ ,  $am$  are perpendicular to  $OD$ .

Now, if the earth had its mean form, a plumb-line at  $A$  would hang in the normal  $zA$  to the mean ellipse; but it hangs actually in  $ZA$ . Hence  $ZAz$  is the deflection (northward in the diagram) which the plumb-line suffers from the local attraction arising from the derangement of the figure and mass of the earth from the mean. This angle is some constant but unknown quantity  $t$ ,  $t$  being reckoned positive when the deflection is northward. This quantity  $t$  is part of the correction  $ZA Z'$ , or  $x$ , added to the observed latitude of  $A$  before applying the principle of least squares. The other part is  $zAZ'$ , which I will now calculate: it is the angle between the two normals drawn through  $A$  to the variable and the mean ellipses. By the property of an ellipse of which the ellipticity is small,

$$ON = 2\epsilon \cdot Om, \text{ and } ON' = 2\epsilon' \cdot Om'.$$

Also as  $Om$ ,  $Om'$ ,  $OD$  differ only by quantities of the order of the ellipticities, they may be put equal to each other in small terms, because we neglect the square of the ellipticities.

$$\begin{aligned} \therefore \angle zAZ' &= \angle NAN' = \angle AN'D - \angle AND \\ &= \tan^{-1} \frac{\cot AND - \cot AN'D}{1 + \cot AND \cot AN'D} = \tan^{-1} \frac{(ND - N'D)AD}{AD^2 + ND \cdot N'D} \\ &= \tan^{-1} \frac{(ON' - ON)AD}{AD^2 + DO^2} = \tan^{-1} \frac{2(\epsilon' - \epsilon)OD \cdot AD}{AD^2 + DO^2} = \tan^{-1}(\epsilon' - \epsilon) \sin 2l \\ &= (\epsilon' - \epsilon) \sin 2l \frac{1''}{\sin 1''}, l \text{ being the observed latitude of } A. \end{aligned}$$

Suppose that  $v$  and  $V$  are the values of  $v$  for the variable and the mean ellipses. Then by the third of the formulæ (1),

$$\angle zAZ' = \frac{\sin 2l}{15000 \sin 1''} (v - V) = 13'' \cdot 75 \sin 2l (v - V) = n(v - V) \text{ suppose. (2).}$$

Hence

$$x = t + n(v - V).$$



Therefore the sum of the squares of errors, which is to be differentiated with respect to  $u$  and  $v$  to obtain a minimum, is

$$\begin{aligned} & (n_1(v-V)+t_1)^2 + (m_1 + \alpha_1 u + \beta_1 v + n_1(v-V) + t_1)^2 \\ & \quad + (m'_1 + \alpha'_1 u + \beta'_1 v + n_1(v-V) + t_1)^2 + \dots \\ & (n_2(v-V)+t_2)^2 + (m_2 + \alpha_2 u + \beta_2 v + n_2(v-V) + t_2)^2 \\ & \quad + (m'_2 + \alpha'_2 u + \beta'_2 v + n_2(v-V) + t_2)^2 + \dots \\ & \quad + \dots = \text{a minimum.} \end{aligned}$$

Let  $U$  and  $V$  be the values of  $u$  and  $v$  which belong to the mean ellipse. These values, then, must be put for  $u$  and  $v$  in the two equations produced by differentiating the above with respect to  $u$  and  $v$ . We have

$$\begin{aligned} & \alpha_1(m_1 + \alpha_1 U + \beta_1 V + t_1) + \alpha'_1(m'_1 + \alpha'_1 U + \beta'_1 V + t_1) + \dots \\ & + \alpha_2(m_2 + \alpha_2 U + \beta_2 V + t_2) + \alpha'_2(m'_2 + \alpha'_2 U + \beta'_2 V + t_2) + \dots \\ & + \dots = 0; \end{aligned}$$

and

$$\begin{aligned} & n_1 t_1 + (\beta_1 + n_1)(m_1 + \alpha_1 U + \beta_1 V + t_1) + (\beta'_1 + n_1)(m'_1 + \alpha'_1 U + \beta'_1 V + t_1) + \dots \\ & n_2 t_2 + (\beta_2 + n_2)(m_2 + \alpha_2 U + \beta_2 V + t_2) + (\beta'_2 + n_2)(m'_2 + \alpha'_2 U + \beta'_2 V + t_2) + \dots \\ & + \dots = 0. \end{aligned}$$

Let  $(m)$  be a symbol representing the sum of all the  $m$ 's appertaining to the divisions of the same Arc; and let  $\Sigma(m)$  represent the sum of all these sums for all the Arcs; and similarly for other quantities besides  $m$ . Then the above equations become

$$\Sigma(m\alpha) + \Sigma(\alpha^2) U + \Sigma(\alpha\beta) V + \Sigma t(\alpha) = 0$$

$$\text{and} \quad \left. \begin{aligned} & \Sigma(m\beta) \} + \Sigma(\alpha\beta) \} U + \Sigma(\beta^2) \} V + \Sigma t(\beta) \} \\ & + \Sigma n(m) \} + \Sigma n(\alpha) \} + \Sigma n(\beta) \} + \Sigma tni \} \end{aligned} \right\} = 0,$$

$i$  being the number of stations on the representative Arc.

The numerical quantities involved in the first two lines of these equations have been already calculated in the article on the Figure of the Earth in the British Ordnance Survey Volume, from which I borrow the results in Table II. on the following page. The quantities involving  $n$  are calculated in Table I., and the results inserted in Table II. with the others.

TABLE I., referred to in last page;  $\log n = \log 13.75 + \log \sin 2 \text{ lat.} = 1.1383027 + \log \sin 2 \text{ lat.}$ 

Arce.	Standard station.	Latitude.	$\log \sin 2 \text{ lat.}$	$\log n$ .	$n$ .	$\log n(m)$ .	$n(m)$ .	$\log n(\alpha)$ .	$n(\alpha)$ .	$\log n(\beta)$ .	$n(\beta)$ .	$z$ .	$ni$ .
Anglo-Gallic.	St. Agnes .....	49° 53' 33".03	1.9936350	1.1319377	13.54995	2.9564540	904.5946	2.6192882	416.1867	2.2251352	-167.9327	34	460.6983
Russian ...	Staro-Nekrassowka }	45° 20' 2.8	1.9999705	1.1382732	13.74907	2.9804517	955.9864	2.8746163	749.2319	2.3419912	-219.7815	13	178.7379
Indian II. ...	Damargida .....	18° 3' 15.29	1.7703483	0.9086510	8.10310	1.1551496	14.2939	1.1082965	-12.8321	1.3254412	-21.1564	8	64.8248
Indian I. ....	Trivandeporum.	11° 44' 52.59	1.6006279	0.7389306	5.48189	1.9403277	0.8716	0.4945768	3.1230	0.4521691	2.8325	2	10.9638
Prussian .....	Trunz .....	54° 13' 11.47	1.9771092	1.1154119	13.04403	1.4679803	29.3752	0.9733472	9.4047	0.4959849	-3.1332	3	39.1321
Peruvian .....	Tarqui, S. ....	3° 4' 32.07	1.0299991	0.1683018	1.47334	1.3296698	0.2136	0.2184108	1.6535	0.2179462	1.6518	2	2.9467
Hanoverian...	Göttingen .....	51° 31' 47.85	1.9886188	1.1269215	13.39435	1.7802305	60.2879	0.9879777	9.7270	0.3962009	-2.4900	2	26.7887
Danish .....	Lauenburg .....	53° 22' 17.05	1.9811875	1.1194902	13.16710	1.2212373	-16.6432	0.8607994	7.2577	0.3519781	-2.2489	2	26.3342
Totals .....	.....	.....	.....	.....	.....	.....	1948.9800	.....	1183.7524	.....	-412.2584		

TABLE II., gathered from the Ordnance Survey Volume, and from Table I.

Arce.	(n).	(α).	(β).	(mα).	(mβ).	(α <sup>2</sup> ).	(αβ).	(β <sup>2</sup> ).	(m).	$\log n(\alpha)$ .	$n(\alpha)$ .	$n(\beta)$ .	$ni$ .
Anglo-Gallic.	+66.760	+30.7150	-12.3936	+118.9207	-45.3653	+155.0671	-35.5339	+11.1906	+904.5946	+416.1867	-167.9327	-167.9327	460.6983
Russian .....	69.531	+54.4933	-15.9852	+386.3623	-126.4488	+335.5318	-112.1421	+29.9745	+955.9864	+749.2319	-219.7815	-219.7815	178.7379
Indian II. ...	1.764	-1.5836	-2.6109	-12.7516	-10.4936	+46.5750	+36.6984	+39.3612	+14.2939	-12.8321	-21.1564	-21.1564	64.8248
Indian I. ....	0.159	+0.5697	+0.5167	+0.0906	+0.0822	+0.3246	+0.2944	+0.2670	+0.8716	+3.1230	+2.8325	+2.8325	10.9638
Prussian .....	2.252	+0.7210	+0.2402	+1.4630	+0.4953	+0.3267	-0.1096	+0.0368	+29.3752	+9.4047	+1.6535	+1.6535	39.1321
Peruvian .....	0.145	+1.1293	+1.1211	+0.1627	+0.1626	+1.2596	+1.2582	+1.2582	+0.2136	+0.2136	+1.6518	+1.6518	2.9467
Hanoverian...	4.501	+0.7262	+0.1859	+3.2686	-0.8367	+0.5274	-0.1350	+0.0346	+60.2879	+9.7270	-2.4900	-2.4900	26.7887
Danish .....	-1.264	+0.5512	-0.1708	-0.6967	+0.2159	+0.3038	-0.0941	+0.0292	-16.6432	+7.2577	-2.2489	-2.2489	26.3342
Totals .....	.....	.....	.....	+496.8196	-183.1790	+539.9160	-109.7637	+82.1508	+1948.9800	+1183.7524	-412.2584	-412.2584	

TABLE III., containing the Logarithms of the numbers in Table II.

Anglo-Galic..	1·8245163	1·4873505	1·0931975	2·0752575	1·6567238	2·1905197	1·5506429	1·0488534	2·9564540	2·6192882	2·2251352	2·6634166
Russian .....	1·8421785	1·7363431	1·2037180	2·5869948	2·1019147	2·5257337	2·0497687	1·6017831	2·9804517	2·8746163	2·3419912	2·2522167
Indian II. ...	0·2464986	0·1996455	0·04167902	1·1055647	1·0209245	1·6681529	1·5646471	1·4677738	1·1551496	1·1082965	1·3254412	1·8117412
Indian I. ....	1·2013971	1·7556462	1·7132385	2·9571282	2·9148718	1·5113485	1·4689378	1·4265113	1·9403172	0·4945768	0·4521691	1·0399611
Prussian .....	0·3525684	1·8579353	1·3805730	0·1652443	1·6948683	1·5141491	1·0398106	2·5658478	1·4679803	0·9733472	0·4959849	1·5925331
Peruvian .....	1·1613680	0·0501090	0·0496444	1·2113876	1·2111205	0·1002327	0·0997497	0·0993007	1·3296698	0·2184108	0·2179462	0·4693359
Hanoverian...	0·6533090	1·8610562	1·2692794	0·5143618	1·9225698	1·7221401	1·1303338	2·5390761	1·7802305	0·9879777	0·3962009	1·4279516
Danish .....	0·1017471	1·7413092	1·2324879	1·8430458	1·3342526	1·4825878	2·9735896	2·4653829	1·2212373	0·8607994	0·3519781	1·4205201
Totals .....	.....	.....	.....	2·6961987	2·2628757	2·7323262	2·0404588	1·9146118	3·2898074	3·0732608	2·6151695	

TABLE IV., derived from Table II.

Arcs.	$(m\beta) + n(m).$	$(\alpha\beta) + n(\alpha).$	$(\beta^2) + n(\beta).$	$(\beta) + m.$	The logarithms of the foregoing numbers.		
Anglo-Galic. ....	859·2293	380·6528	— 156·7421	448·3047	2·9341091	2·5805290	2·6515733
Russian .....	829·5376	637·0898	— 179·8070	162·7527	2·9188361	2·8042007	2·2115282
Indian II. ....	3·8003	23·8663	8·2048	62·2139	0·5798179	1·3777851	1·7938874
Indian I. ....	0·9538	3·4174	3·0995	11·4805	1·9794573	0·5336958	1·0599608
Prussian .....	28·8799	9·2951	— 3·0964	38·8919	1·4605957	0·9682541	1·5898592
Peruvian .....	0·3762	2·9117	2·9087	4·0678	1·5754188	0·4641466	0·6093596
Hanoverian .....	59·4512	9·5920	— 2·4554	26·6028	1·7741607	0·9819092	1·4249273
Danish .....	— 16·4273	7·1636	— 2·2197	26·1634	— 1·2155662	0·8551313	1·4176941
Totals .....	1765·8010	1073·9887	— 330·1076	.....	3·2469417	— 2·5186555	

8. I will now apply the formulæ just obtained to determine the Mean Figure of the Earth from the data afforded by the eight arcs. For convenience I shall use the well-known symbol ( $2\cdot6961987$ ) to mean the number of which  $2\cdot6961987$  is the logarithm; and so of other numbers. By substitution from the Table, the formulæ give

$$\begin{aligned} & (2\cdot6961987) + (2\cdot7323262)U - (2\cdot0404588)V + (1\cdot4873505)t_1 \\ & + (1\cdot7363431)t_2 - (0\cdot1996455)t_3 + (\bar{1}\cdot7556462)t_4 + (\bar{1}\cdot8579353)t_5 \\ & + (0\cdot0501090)t_6 + (\bar{1}\cdot8610562)t_7 + (\bar{1}\cdot7413092)t_8 = 0, \\ & (3\cdot2469417) + (3\cdot0309997)U - (2\cdot5186555)V + (2\cdot6515733)t_1 \\ & + (2\cdot2115282)t_2 + (1\cdot7938874)t_3 + (1\cdot0599608)t_4 + (1\cdot5898592)t_5 \\ & + (0\cdot6093596)t_6 + (1\cdot4249273)t_7 + (1\cdot4176941)t_8 = 0. \end{aligned}$$

Multiplying by the coefficients of V crosswise, and subtracting so as to eliminate V, we have

$$\begin{aligned} & (5\cdot2148542) + (5\cdot2509817)U + (4\cdot0060060)t_1 + (4\cdot2549986)t_2 \\ & - (2\cdot7183010)t_3 \\ & - (5\cdot2874005) - (5\cdot0714585)U - (4\cdot6920321)t_1 - (4\cdot2519870)t_2 \\ & - (3\cdot834362)t_3 \\ & + (2\cdot2743017)t_4 + (2\cdot2765908)t_5 + (2\cdot5687645)t_6 + (2\cdot3797117)t_7 \\ & + (2\cdot2599647)t_8 \\ & - (3\cdot1004196)t_4 - (3\cdot6303180)t_5 - (2\cdot6498184)t_6 - (3\cdot4653861)t_7 \\ & - (3\cdot4581529)t_8 = 0. \end{aligned}$$

Putting numbers in the place of logarithms,

$$\begin{aligned} & 164004 + 178230U + 10139t_1 + 17989t_2 - 523t_3 \\ & - 193821 - 117885U - 49208t_1 - 17864t_2 - 6829t_3 \\ & - 29817 + 60345U - 39069t_1 + 125t_2 - 7352t_3 \\ & + 188t_4 + 238t_5 + 370t_6 + 240t_7 + 182t_8 \\ & - 1260t_4 - 4269t_5 - 446t_6 - 2920t_7 - 2872t_8 \\ & - 1072t_4 - 4031t_5 - 76t_6 - 2680t_7 - 2690t_8 = 0. \end{aligned}$$

Putting logarithms in the place of numbers,

$$\begin{aligned} & - (4\cdot4744639) + (4\cdot7806413)U - (4\cdot5918323)t_1 + (2\cdot0969100)t_2 \\ & - (3\cdot8664055)t_3 - (3\cdot0301948)t_4 - (3\cdot6054128)t_5 - (1\cdot8808136)t_6 \\ & - (3\cdot4281348)t_7 - (3\cdot4297523)t_8 = 0. \end{aligned}$$

Transposing and dividing by the coefficient of U,

$$\begin{aligned} U = & (\bar{1}\cdot6938226) + (\bar{1}\cdot8111910)t_1 - (\bar{3}\cdot3162687)t_2 + (\bar{1}\cdot0857642)t_3 \\ & + (\bar{2}\cdot2495535)t_4 + (\bar{2}\cdot8247715)t_5 + (\bar{3}\cdot1001723)t_6 + (\bar{2}\cdot6474933)t_7 \\ & + (\bar{2}\cdot6491110)t_8. \end{aligned}$$

Now 2089 = (3·3199384),

$$\begin{aligned}\therefore 2089U &= (3\cdot0137610) + (3\cdot1311294)t_1 - (0\cdot6362071)t_2 + (2\cdot4057026)t_3 \\ &\quad + (1\cdot5694919)t_4 + (2\cdot1447099)t_5 + (0\cdot4201107)t_6 \\ &\quad + (1\cdot9674319)t_7 + (1\cdot9690494)t_8 \\ &= 1032\cdot2 + 1352\cdot4t_1 - 4\cdot3t_2 + 254\cdot5t_3 + 37\cdot1t_4 + 139\cdot5t_5 + 2\cdot6t_6 \\ &\quad + 92\cdot8t_7 + 93\cdot1t_8.\end{aligned}$$

Transposing the term in V in the first equation of this paragraph and dividing by its coefficient, we have

$$\begin{aligned}V &= (0\cdot6557399) + (0\cdot6918674)U + (\bar{1}\cdot4468917)t_1 + (\bar{1}\cdot6958843)t_2 \\ &\quad - (\bar{2}\cdot1591867)t_3 + (\bar{3}\cdot7151874)t_4 + (\bar{3}\cdot8174765)t_5 + (\bar{2}\cdot0096502)t_6 \\ &\quad + (\bar{3}\cdot8205974)t_7 + (\bar{3}\cdot7008504)t_8 \\ &= (0\cdot6557399) + (\bar{1}\cdot4468917)t_1 + (\bar{1}\cdot6958843)t_2 - (\bar{2}\cdot15918\cdot7)t_3 + (\bar{3}\cdot7151874)t_4 \\ &\quad + (0\cdot3856900) + (0\cdot5030584)t_1 - (\bar{2}\cdot0081361)t_2 + (\bar{1}\cdot777\cdot3316)t_3 + (\bar{2}\cdot9414209)t_4 \\ &\quad + (\bar{3}\cdot8174765)t_5 + (\bar{2}\cdot0096502)t_6 + (\bar{3}\cdot8205974)t_7 + (\bar{3}\cdot7008504)t_8 \\ &\quad + (\bar{1}\cdot5166389)t_5 + (\bar{3}\cdot7920397)t_6 + (\bar{1}\cdot3393609)t_7 + (\bar{1}\cdot3409784)t_8.\end{aligned}$$

Now 417800 = (5·6209684),

$\therefore$  417800V

$$\begin{aligned}&= (6\cdot2767083) + (5\cdot0678601)t_1 + (5\cdot3168527)t_2 - (3\cdot7801551)t_3 + (3\cdot3361558)t_4 \\ &\quad + (6\cdot0066584) + (6\cdot1240268)t_1 - (3\cdot6291045)t_2 + (5\cdot3986000)t_3 + (4\cdot5623893)t_4 \\ &\quad + (3\cdot4384449)t_5 + (3\cdot6306186)t_6 + (3\cdot4415658)t_7 + (3\cdot3218188)t_8 \\ &\quad + (5\cdot1376073)t_5 + (3\cdot4130081)t_6 + (4\cdot9602293)t_7 + (4\cdot9619468)t_8 \\ &= 1891073 + 116912t_1 + 207421t_2 - 6028t_3 + 2168t_4 \\ &\quad + 1015450 + 1330537t_1 - 4257t_2 + 250380t_3 + 36508t_4 \\ &\quad + 2906523 + 1447449t_1 + 203164t_2 + 244352t_3 + 38676t_4 \\ &\quad + 2744t_5 + 4272t_6 + 2764t_7 + 2098t_8 \\ &\quad + 137280t_5 + 2588t_6 + 91249t_7 + 91611t_8 \\ &\quad + 142024t_5 + 6860t_6 + 94013t_7 + 93709t_8.\end{aligned}$$

Substituting the values of 2089 U and 417800 V above deduced in the formulæ (1) of paragraph 6, we have

$$\begin{aligned}\frac{a+b}{2} &= 20888968 - 1352\cdot4t_1 + 4\cdot3t_2 - 254\cdot5t_3 - 37\cdot1t_4 - 139\cdot5t_5 - 2\cdot6t_6 \\ &\quad - 92\cdot8t_7 - 93\cdot1t_8, \\ \frac{a-b}{2} &= \frac{1}{600} \left\{ \frac{a+b}{2} + 417800 V \right\} \\ &= \frac{1}{600} \{ 23795491 + 1446097t_1 + 203168t_2 + 244098t_3 + 38639t_4 \\ &\quad + 139884t_5 + 6857t_6 + 93920t_7 + 93616t_8 \} \\ &= 39659 + 2410\cdot2t_1 + 338\cdot6t_2 + 406\cdot8t_3 + 64\cdot4t_4 + 233\cdot1t_5 + 11\cdot4t_6 \\ &\quad + 156\cdot5t_7 + 156\cdot0t_8;\end{aligned}$$

$$\begin{aligned} \therefore a &= 20928627 + 1057.8t_1 + 342.9t_2 + 152.3t_3 + 27.3t_4 + 93.6t_5 + 8.8t_6 \\ &\quad + 63.7t_7 + 62.9t_8, \\ b &= 20849309 - 3762.6t_1 - 334.3t_2 - 661.3t_3 - 101.5t_4 - 372.6t_5 \\ &\quad - 14.0t_6 - 249.3t_7 - 249.1t_8. \end{aligned}$$

From these we may easily deduce the ellipticity

$$\epsilon = \frac{1}{263.9} \{ 1 + 0.0608t_1 + 0.0085t_2 + 0.0103t_3 + 0.0016t_4 + 0.0059t_5 \\ + 0.0003t_6 + 0.0039t_7 + 0.001639t_8 \}.$$

These formulæ for the semiaxes and ellipticity of the mean figure of the earth show us that the effect of local attraction upon the final numerical results may be very considerable: for example, a deflection of the plumb-line of only 5'' at the standard station (St. Agnes) of the Anglo-Gallic arc would introduce a correction of about one mile to the length of the semi-major axis, and more than three miles to the semi-minor axis. If the deflection at the standard station (Damargida) of the Indian Great Arc be what the mountains and ocean make it (without allowing any compensating effect from variations in density in the crust below, which no doubt exist, but which are altogether unknown), viz. about 17''·24, the semiaxes will be subject to a correction, arising from this cause alone, of half a mile and two miles. This is sufficient to show how great a degree of uncertainty local attraction, if not allowed for, introduces into the determination of the mean figure. As long as we have no means of ascertaining the amount of local attraction at the several standard-stations of the arcs employed in the calculation, this uncertainty regarding the mean figure, as determined by geodesy, must remain.

§ 3. *Comparison of the Anglo-Gallic, Russian, and Indian Arcs, with a view to deduce the Mean Figure of the Earth.*

9. The first three of the eight arcs which have been used in the calculation, viz. the Anglo-Gallic, Russian, and Indian, are of considerable length; and as the *à priori* probability appears to be that the earth nowhere departs much from its mean form, it seems not unlikely that by the following device we may overcome the difficulty pointed out in the last paragraph. I will deduce expressions for the semiaxes of the mean figure of each of these three arcs by the method there given. If reasonable values can be assigned to the expressions for the deflection of the plumb-line from the normals to these three ellipses such as will make the axes the same, we shall have a very strong argument in favour of those being the actual deflections in nature, and of the figure thus deduced, as common to the three arcs, being in fact the mean figure of the earth.

10. In the previous calculation  $t$  has represented the angle which the plumb-line makes, in the plane of the meridian, with the normal to the mean ellipse of the earth. I shall now use  $T$  as the angle which the plumb-line makes, in the plane of the meridian, with the normal to the mean

ellipse of the particular arc under consideration. I shall begin with the Anglo-Gallic arc. Proceeding precisely as in paragraph 8, we have

$$\begin{aligned}(2\cdot0752575) + (2\cdot1905197)U_1 - (1\cdot5506429)V_1 + (1\cdot4873505)T_1 &= 0, \\ (2\cdot9341091) + (2\cdot5805290)U_1 - (2\cdot1951856)V_1 + (2\cdot6515733)T_1 &= 0, \\ (4\cdot2704431) + (4\cdot3857053)U_1 + (3\cdot6825361)T_1 \\ - (4\cdot4847520) - (4\cdot1311719)U_1 - (4\cdot2022162)T_1 &= 0, \\ 18640 + 24306 U_1 + 4814 T_1 \\ - 30532 - 13526 U_1 - 15930 T_1 \\ - 11892 + 10780 U_1 - 11116 T_1 &= 0,\end{aligned}$$

or

$$\begin{aligned}-(4\cdot0752549) + (4\cdot0326188)U_1 - (4\cdot0459485)T_1 &= 0; \\ \therefore U_1 &= (0\cdot0426361) + (0\cdot0133297)T_1, \quad 2089 = (3\cdot3199384), \\ 2089U_1 &= (3\cdot3625745) + (3\cdot3332681)T_1 = 2304\cdot5 + 2154\cdot1 T_1.\end{aligned}$$

By the first of the equations in  $V_1$ , we have

$$\begin{aligned}V_1 &= (0\cdot5246146) + (0\cdot6398768)U_1 + (\bar{1}\cdot9367076)T_1 \\ &= (0\cdot5246146) + (\bar{1}\cdot9367076)T_1 \\ &\quad + (0\cdot6825129) + (0\cdot6532065)T_1, \quad 417800 = (5\cdot6209684); \\ \therefore 417800V_1 &= (6\cdot1455830) + (6\cdot3034813) + \{ (5\cdot5576760) + (6\cdot2741749) \} T_1 \\ &= 1398244 + 2011320 + \{ 361140 + 1880074 \} T_1 \\ &= 3409564 + 2241214 T_1; \\ \therefore \frac{a_1 + b_1}{2} &= 20887695 - 2154\cdot1 T_1 \\ \frac{a_1 - b_1}{2} &= \frac{1}{600} \{ 24297259 + 2239060 T_1 \} = 40495 + 3731\cdot8 T_1; \\ \therefore a_1 &= 20928190 + 1577\cdot7 T_1, \quad b_1 = 20847200 - 5885\cdot9 T_1, \\ \epsilon_1 &= \frac{1}{258\cdot4} (1 + 0\cdot0921 T_1).\end{aligned}$$

11. I proceed to the second, the Russian arc.

$$\begin{aligned}(2\cdot5869948) + (2\cdot5257337)U_2 - (2\cdot0497688)V_2 + (1\cdot7363431)T_2 &= 0, \\ (2\cdot9188361) + (2\cdot8042007)U_2 - (2\cdot2548066)V_2 + (2\cdot2115282)T_2 &= 0, \\ (4\cdot8418014) + (4\cdot7805403)U_2 + (3\cdot9911497)T_2 \\ - (4\cdot9686049) - (4\cdot8539695)U_2 - (4\cdot2612970)T_2 &= 0, \\ 69471 + 60331 U_2 + 9798 T_2 \\ - 93026 - 71445 U_2 - 18251 T_2 \\ - 23555 - 11114 U_2 - 8453 T_2 &= 0,\end{aligned}$$

or

$$\begin{aligned}-(4\cdot3720831) - (4\cdot0458704)U_2 - (3\cdot9270109)T_2 &= 0; \\ \therefore U_2 &= -(0\cdot3262127) - (\bar{1}\cdot8811405)T_2, \quad 2089 = (3\cdot3199384), \\ 2089U_2 &= -(3\cdot6461511) - (3\cdot2010789)T_2 = -4427\cdot4 - 1588\cdot8 T_2.\end{aligned}$$

By the first of the equations in  $V_2$ , we have

$$\begin{aligned} V_2 &= (0.5372260) + (0.4759649)U_2 + (\bar{1}.6865743)T_2 \\ &= (0.5372260) - (0.8021776) + \{( \bar{1}.6865743) - (0.3571054)\}T_2 \\ 417800 &= (5.6209684); \end{aligned}$$

$$\begin{aligned} \therefore 417800 V_2 &= (6.1581944) - (6.4231460) + \{(5.3075427) - (5.9780738)\}T_2 \\ &= 1439443 - 2649391 + \{203022 - 950766\}T_2 \\ &= -1209948 - 747744 T_2; \end{aligned}$$

$$\therefore \frac{a_2 + b_2}{2} = 20894427 + 1588.8 T_2,$$

$$\frac{a_2 - b_2}{2} = \frac{1}{600} \{19684479 - 746155 T_2\} = 32807 - 1243.6 T_2,$$

$$a_2 = 20927234 + 345.2 T_2, \quad b_2 = 20861620 + 2832.4 T_2,$$

$$\epsilon = \frac{1}{318.9} (1 - 0.0379 T_2).$$

12. The following is the calculation for the Indian arc:—

$$\begin{aligned} -(1.1055647) + (1.6681529)U_3 + (1.5646471)V_3 - (0.1996455)T_3 &= 0, \\ (0.5798179) + (1.3777851)U_3 + (0.9140680)V_3 + (1.7938874)T_3 &= 0, \end{aligned}$$

$$-(2.0196327) + (2.5822209)U_3 - (1.1137135)T_3$$

$$-(2.1444650) - (2.9424322)U_3 - (3.3585345)T_3 = 0,$$

$$-105 + 382U_3 - 13 T_3$$

$$-139 - 876 U_3 - 2283 T_3$$

$$-244 - 494 U_3 - 2296 T_3 = 0,$$

or

$$-(2.3873898) - (2.6937269)U_3 - (3.3609719)T_3 = 0;$$

$$\therefore U_3 = -(\bar{1}.6936629) - (0.6672450)T_3, \quad 2089 = (3.3199384),$$

$$2089 U_3 = -(3.0136013) - (3.9871834)T_3 = -1031.8 - 9709.2 T_3.$$

By the first of the equations in  $V_3$ , we have

$$\begin{aligned} V_3 &= (\bar{1}.5409176) - (\bar{0}.1035058)U_3 + (\bar{2}.6349984)T_3 \\ &= (\bar{1}.5409176) + (\bar{1}.7971687) + \{\bar{2}.6349984\} + (0.7707508)\}T_3 \\ 417800 &= (5.6209684), \end{aligned}$$

$$\begin{aligned} 417800 V_3 &= (5.1618860) + (5.4181371) + \{(4.2559668) + (6.3917192)\}T_3 \\ &= 145173 + 261901 + \{18029 + 2464445\}T_3 \\ &= 407074 + 2482474 T_3; \end{aligned}$$



$$\therefore \frac{a_3 + b_3}{2} = 20891032 + 9709 \cdot 2 T_3,$$

$$\frac{a_3 - b_3}{2} = \frac{1}{600} \{21298106 + 2492183 T_3\} = 35497 + 4153 \cdot 6 T_3;$$

$$\therefore a_3 = 20926529 + 13862 \cdot 8 T_3, \quad b_3 = 20855535 + 5555 \cdot 6 T_3,$$

$$e_3 = \frac{1}{294 \cdot 8} (1 + 0 \cdot 1170 T_3).$$

13. I have now, if possible, to find values of  $T_1$ ,  $T_2$ ,  $T_3$  which will make these three ellipses, which measure the Anglo-Gallic, the Russian, and the Indian arcs, the same; that is,  $a_1 = a_2 = a_3$ ,  $b_1 = b_2 = b_3$ . These give the four following equations:—

$$\left. \begin{aligned} 1577 \cdot 7 T_1 - 345 \cdot 2 T_2 + 956 &= 0, \\ 5885 \cdot 9 T_1 + 2832 \cdot 4 T_2 + 14420 &= 0, \\ 1577 \cdot 7 T_1 - 13862 \cdot 8 T_3 + 1661 &= 0, \\ 5885 \cdot 9 T_1 + 5555 \cdot 6 T_3 + 8335 &= 0, \end{aligned} \right\} \text{or} \left\{ \begin{aligned} (3 \cdot 1980244) T_1 - (2 \cdot 5380708) T_2 + (2 \cdot 9804579) &= 0, \\ (3 \cdot 7698129) T_1 + (3 \cdot 4521546) T_2 + (4 \cdot 1589653) &= 0, \\ (3 \cdot 1980244) T_1 - (4 \cdot 1418509) T_3 + (3 \cdot 2203696) &= 0, \\ (3 \cdot 7698129) T_1 + (3 \cdot 7447310) T_3 + (3 \cdot 9209056) &= 0, \end{aligned} \right.$$

The most likely solutions of these four equations connecting the three quantities  $T_1$ ,  $T_2$ ,  $T_3$  which we are seeking are found by the method of least squares. This leads to the three following equations:—

$$\left. \begin{aligned} 2(6 \cdot 3960488) \} T_1 - (5 \cdot 7360952) \} T_2 - (7 \cdot 3398753) \} T_3 + (6 \cdot 1784823) + (6 \cdot 4183940) \\ + 2(7 \cdot 5396258) \} + (7 \cdot 2219675) \} + (7 \cdot 5145439) \} + (7 \cdot 9287782) + (7 \cdot 6907185) &= 0. \end{aligned} \right.$$

$$\left. \begin{aligned} (5 \cdot 7360952) \} T_1 - (5 \cdot 0761416) \} T_2 + (5 \cdot 5185287) \\ + (7 \cdot 2219675) \} + (6 \cdot 9043092) \} + (7 \cdot 6111199) &= 0, \end{aligned} \right.$$

$$\left. \begin{aligned} (7 \cdot 3398753) \} T_1 - (8 \cdot 2837018) \} T_3 + (7 \cdot 3622205) \\ + (7 \cdot 5145439) \} + (7 \cdot 4894620) \} + (7 \cdot 6656366) &= 0, \end{aligned} \right.$$

or

2 × 2489	— 545		— 21871	+ 1508	+ 2621
+ 2 × 34644	+ 16671		+ 32700	+ 84875	+ 49059
74266 $T_1$ + 16126 $T_2$			+ 10829 $T_3$ + 138063 = 0		
(4·8707900) (4·2075267)			(4·0345884) (5·1400773)		
545	— 119	+ 330	21871	— 192177	+ 23026
+ 16671	+ 8022	+ 40843	+ 32700	+ 30865	+ 46306
17216 $T_1$ + 7903 $T_2$ + 41173 = 0,			54571 $T_1$ — 161312 $T_3$ + 69332 = 0		
(4·2359323) (3·8977920) (4·6146125)			(4·7369619) (5·2076667) (4·8409337);		

$$\therefore T_2 = -(0 \cdot 3381403) T_1 - (0 \cdot 7168205), \quad T_3 = (\bar{1} \cdot 5292952) T_1 + (1 \cdot 6332670);$$

$$\therefore 16126 T_2 = -(4 \cdot 5456670) T_1 - (4 \cdot 9243472), \quad 10829 T_3 = (3 \cdot 5638836) T_1 + (3 \cdot 6678554)$$

$$= -35129 T_1 - 84013, \quad = 3663 T_1 + 4654;$$

$$\therefore \{74266 - 35129 + 3663\} T_1 + 138063 - 84013 + 4654 = 0,$$

$$42800 T_1 + 58704 = 0, \quad T_1 = -1'' \cdot 37,$$

$$T_2 = -2 \cdot 18 T_1 - 5 \cdot 21 = +2 \cdot 99 - 5 \cdot 21 = -2'' \cdot 22,$$

$$T_3 = 0 \cdot 338 T_1 + 0 \cdot 430 = -0 \cdot 463 + 0 \cdot 430 = -0'' \cdot 033.$$

When these are substituted in the semiaxes, they give

$$\begin{aligned} a_1 &= 20928190 - 2161 = 20926029, & a_2 &= 20927234 - 766 = 20926468, \\ a_3 &= 20926529 - 457 = 20926072, \\ b_1 &= 20847200 + 8064 = 20855264, & b_2 &= 20861620 - 6288 = 20855332, \\ b_3 &= 20855535 - 183 = 20855352. \end{aligned}$$

These three results are remarkably near each other; they differ from their average, 20926189 and 20855316, in no case by so much as 300 feet, and in most cases by much less. I think, then, that we may safely infer that this average ellipse is in fact the mean figure of the earth. This being the case,  $T_1, T_2, T_3$  are the same as  $t_1, t_2, t_3$ ; and therefore the deflections of the plumb-line in the meridian at the standard stations of the Anglo-Gallic, Russian, and Indian arcs are  $1''\cdot37, 2''\cdot22, 0''\cdot033$ , all in the southern direction\*.

14. The values, then, which I would assign to the semiaxes and ellipticity of the Mean Figure of the Earth are as follows:—

$$a = 20926180, \quad b = 20855316 \text{ feet}, \quad e = \frac{1}{295\cdot3}.$$

If these are substituted in the formulæ (1) of paragraph (6), we have

$$U = -0\cdot3581 \text{ and } V = 0\cdot8819.$$

#### § 4. *Speculations regarding the constitution of the Earth's Crust.*

15. If the reasoning in the last section, which has led to so satisfactory a result, be correct, I think we may draw some useful inferences regarding the constitution of the earth's crust.

By substituting the values of  $U, V, t_1, t_2, t_3$  in the formulæ similar to  $m + \alpha U + \beta V + t$  for the fifty-five stations of the eight arcs, which will be found at p. 766 of the Ordnance Survey Volume, every one of the results will be small. These results are the corrections of the latitudes of the stations in referring them to the mean ellipse; that is, they are the deflections of the plumb-line in the meridian at those stations owing to local attraction, or the attraction arising from the departure of the actual figure of the earth from the mean figure.

Fifteen of these formulæ I here select, adding one new one for Dehra about 56 miles to the north of Kaliana, the northern extremity of the Indian arc. They are as follows:—

\* The numerical calculations in paragraphs 7 to 13 inclusive have been tested at the Government Trigonometrical Survey Office in Calcutta.

*From the Anglo-Gallic Arc.*

		Deflec- tions.	Calculated attractions.	Deflections to be ac- counted for.
(1) Barcelona .....	+1.440-3.0644 U+0.0553 V-1.37=	+2.22		
(2) Dunkirk .....	+0.767+0.4115 U-0.0765 V-1.37=	-0.84		
(3) High Port Cliff.	+1.778+0.2532 U-0.0450 V-1.37=	+1.28	+3.29	-2.01
(4) Week Down ...	+1.747+0.2539 U-0.0452 V-1.37=	+0.25	+1.98	-1.73
(5) Boniface Down	+1.967+0.2559 U-0.0455 V-1.37=	+0.46	+2.42	-1.96
(6) Dunnose .....	-0.499+0.2613 U-0.0466 V-1.37=	-2.00	-0.54	-1.46
(7) Blackdown ...	+4.279+0.2859 U-0.0513 V-1.37=	+2.76		
(8) Burleigh Moor.	-1.814+1.6845 U-0.4137 V-1.37=	-4.15	-4.55	+0.40
(9) Cowhythe .....	-6.915+2.8048 U-0.8340 V-1.37=	-9.31	-5.50?	-3.81
(10) Ben Hutig .....	+0.095+3.1173 U-0.9708 V-1.37=	-3.25	-2.01	-1.24
(11) Saxavord .....	+4.403+3.9370 U-1.3699 V-1.37=	+0.41		

*From the Russian Arc.*

(12) Tornea .....	+11.826+7.3799 U-2.5821 V-2.22=	+4.69		
(13) Fuglences .....	+10.008+9.1231 U-3.8418 V-2.22=	+1.13		

*From the Indian Arc.*

(14) Punnce .....	+ 0.625-3.5622U-3.1853V-0.033=	-0.94	+22.71	-23.65
(15) Kaliaua .....	+ 0.403+4.1251U+2.7756V-0.033=	+1.34	+34.16	-32.82
(16) Dehra* .....	+53.796+4.4215U-0.1010V-0.033=	+52.09		

I have inserted the formula of Cowhythe from p. 771 of the Ordnance Survey Volume. I have also added two columns, in one of which are given the deflections of the plumb-line arising from attraction at those of the stations for which it has been calculated. For those of the Anglo-Gallic Arc, I refer to the Ordnance Survey Volume, sect. xi. p. 625; and for those of the Indian Arc to my paper in the Philosophical Transactions for 1861, p. 593. I would observe that not only in the two stations of the Indian Arc, but in those I have selected from the Anglo-Gallic Arc (all of which are near the sea-shore), allowance is made for deficiency of density and attraction of sea-water. In the stations (3), (4), (5), (6) the effect of the sea for about 9 miles south of the coast is taken and estimated at  $+0''.27$  (see Ordnance Survey Volume, p. 631); in station (8) for 36 miles north, and estimated at  $-0''.39$  (p. 642); in station (9) for 50 miles north, and estimated at  $-0''.70$  (p. 664); in station (10) for 50 miles north, and estimated at  $-0''.64$  (p. 662). It is of importance to bear this in mind. For stations (14) and (15) the effect of the sea the whole way to the south pole

\* This is calculated by the formulæ at p. 737 of the Ordnance Survey Volume, from the following data obligingly furnished me by Major Walker, Superintendent of the Government Trigonometrical Survey of India, viz.

Astronomical latitude of Dehra  $30^{\circ} 19' 19''$ .

Distance of parallels of Dehra and Damargida 4463510.7 feet.

The latitude of Damargida is  $18^{\circ} 3' 15''$ .

is taken, and estimated at  $+19''\cdot71$  and  $+6''\cdot18$ , the effect of the mountain mass on the north being  $+3''\cdot00$  and  $27''\cdot98$ .

16. The first thing I observe in the results given in the last paragraph is the very small amount of the resultant deflections at the two extremities of the Indian Arc—Punnœ close to Cape Comorin, and Kaliaua the nearest station to the Himmalaya Mountains; whereas the effect of the Ocean and the Mountains has been shown to be very large. This shows that the effect of variations of density in the crust must be very great, in order to bring about this near compensation. In fact the density of the crust beneath the mountains must be less than that below the plains, and still less than that below the ocean-bed. If solidification from the fluid state commenced at the surface, the amount of contraction in the solid parts beneath the mountain-region has been less than in the parts beneath the sea. In fact, it is this unequal contraction which appears to have caused the hollows in the external surface which have become the basins into which the waters have flowed to form the ocean. As the waters flowed into the hollows thus created, the pressure on the ocean-bed would be increased, and the crust, so long as it was sufficiently thin to be influenced by hydrostatic principles of floatation, would so adjust itself that the pressure on any *couche de niveau* of the fluid should remain the same. At the time that the crust first became sufficiently thick to resist fracture under the strain produced by a change in its density—that is, when it first ceased to depend for the elevation or depression of its several parts upon the principles of floatation, the total amount of matter in any vertical prism, drawn down into the fluid below to a given distance from the earth's centre, had been the same through all the previous changes. After this, any further contraction or any expansion in the solid crust would not alter the amount of matter in the vertical prism, except where there was an ocean; in the case of greater contraction under an ocean than elsewhere, the ocean would become deeper and the amount of matter greater, and in case of a less contraction or of an expansion of the crust under an ocean, the ocean would become shallower, or the amount of matter in the vertical prism less than before. It is not likely that expansion and contraction in the solid crust would affect the arrangement of matter in any other way. That changes of level do take place, by the rising and sinking of the surface, is a well-established fact, which rather favours these theoretical considerations. But they receive, I think, great support from the other fact, that the large effect of the ocean at Punnœ and of the mountains at Kaliaua almost entirely disappear from the resultant deflections brought out by the calculations. The formulæ of paragraph 15 show that when we get close to the mountain-mass, as at Dehra, which is at the foot of the mountains where they first rise rapidly above the plains, the resultant deflection is very great; the less density of the crust down below the sea-level drawn under the mountain-mass has here a very trifling influence. This is as it should be, if the depth of this less density is considerable;

whereas at Kaliana, and stations still further off, the attraction of the mountain-mass above the sea-level, and the deficiency of attraction from the crust below that level, would nearly counterbalance each other. Thus, if the thickness of the crust below the plains is 100 miles, and the amount of matter in the crust under the plains equals that of the crust and mountains together in the mountain-region, then the deflections at Kaliana, Kalianpur, and Damargida, instead of being  $27''.98$ ,  $12''.05$ ,  $6''.79$ , arising from the mountains alone, are reduced to  $1''.54$ ,  $-0''.06$ ,  $-0''.06$  (see *Philosophical Transactions* for 1858, p. 759), which are all insignificant compared with the large deflections caused by the mountains alone.

This theory, that the wide ocean has been collected on parts of the earth's surface where hollows have been made by the contraction and therefore increased density of the crust below, is well illustrated by the existence of a whole hemisphere of water, of which New Zealand is the pole, in stable equilibrium. Were the crust beneath only of the same density as that beneath the surrounding continents, the water would be drawn off by attraction and not allowed to stand in the undisturbed position it now occupies.

17. I have, in what goes before, supposed that, in solidifying, the crust contracts and grows denser, as this appears to be most natural, though, after the solid mass is formed, it may either expand or contract, according as an accession or diminution of heat may take place. If, however, in the process of solidifying, the mass becomes lighter, the same conclusion will follow—the mountains being formed by a greater degree of expansion of the crust beneath them, and not by a less contraction, than in the other parts of the crust. It may seem at first difficult to conceive how a crust could be formed at all, if in the act of solidification it becomes heavier than the fluid on which it rests; for the equilibrium of the heavy crust floating on a lighter fluid would be unstable, and the crust would sooner or later be broken through, and would sink down into the fluid, which would overflow it. If, however, this process went on perpetually, the descending crust, which was originally formed by a loss of heat radiated from the surface into space, would reduce the heat of the fluid into which it sank, and after a time a thicker crust would be formed than before, and the difficulty of its being broken through would become greater every time a new one was formed. Perhaps the tremendous dislocation of stratified rocks in huge masses with which a traveller in the mountains, especially in the interior of the Himalaya region, is familiar, may have been brought about in this way. The catastrophes, too, which geology seems to teach have at certain epochs destroyed whole species of living creatures, may have been thus caused, at the same time breaking up the strata in which those species had for ages before been deposited as the strata were formed. These phenomena must now long have ceased to occur, at any rate on a very extensive scale, as Mr. Hopkins's investigations on Precession appear

to prove that the crust is very thick, at least 800 or 1000 miles; and this result, I understand, has been recently confirmed by Professor W. Thomson in a paper "On the Rigidity of the Earth."

18. These theoretical considerations receive, I think, some confirmation from an examination of the calculated deflection of the plumb-line at stations near the sea-shore. It is for this reason that I have collected the thirteen examples from the Anglo-Gallic and Russian Arcs in paragraph 15, all of which are near the coast. The evidence they furnish, however, is not to be compared in weight with that of the Indian Arc, already considered. In some instances the local attraction of the surrounding country and of the ocean for a certain distance has been calculated, as already stated. These results I will take into account, except the allowances for the ocean as noted at the end of paragraph 15, which I deduct in the following arrangement of the stations.

*The Stations at which the Deflection is towards the Land.*

	Deflection.
(1) Barcelona, lat. $41^{\circ} 23'$ , S.E. coast of Spain, .....	$+2^{\circ} 22'$
(2) Dunkirk, ,, $51^{\circ} 2'$ , N.N.W. France, .....	$-0^{\circ} 84'$
(7) Blackdown, ,, $50^{\circ} 41'$ , S. ,, Dorset, .....	$+2^{\circ} 76'$
(9) Cowhythe, ,, $57^{\circ} 41'$ , N. ,, Banff, .....	$-3^{\circ} 81' + 0^{\circ} 70' = -3^{\circ} 11'$
(10) BenHutig, ,, $58^{\circ} 33'$ , N. ,, Sutherland, .....	$-1^{\circ} 24' + 0^{\circ} 64' = -0^{\circ} 60'$
(12) Tornea, ,, $65^{\circ} 50'$ , S. ,, Lapland, .....	$+4^{\circ} 69'$

*The Stations at which the Deflection is towards the Sea.*

	Deflection.
(3) High Port Cliff, $50^{\circ} 36'$ , S. coast of Isle of Wight, .....	$-2^{\circ} 01' - 0^{\circ} 27' = -2^{\circ} 28'$
(4) Week Down, $50^{\circ} 36'$ , ,, ,, .....	$-1^{\circ} 73' - 0^{\circ} 27' = -2^{\circ} 00'$
(5) Boniface Down, $50^{\circ} 36'$ , ,, ,, .....	$-1^{\circ} 96' - 0^{\circ} 27' = -2^{\circ} 23'$
(6) Dunnose, $50^{\circ} 37'$ , ,, ,, .....	$-1^{\circ} 46' - 0^{\circ} 27' = -1^{\circ} 73'$
(8) Burleigh Moor, $54^{\circ} 34'$ , N. coast of Yorkshire, .....	$+0^{\circ} 40' + 0^{\circ} 39' = +0^{\circ} 79'$
(11) Saxavord, $60^{\circ} 50'$ , N. ,, Unst, .....	$+0^{\circ} 41'$
(13) Fuglencæs, $70^{\circ} 40'$ , N. ,, Finmark, .....	$+1^{\circ} 13'$

The theory I have proposed, that contraction of the crust has formed the basins in which the sea has settled, can hardly be expected to apply so completely to such confined sheets of water as the Mediterranean south of Spain, and the Gulf of Bothnia. Here there may be an actual deficiency of attracting matter in the water, not altogether compensated for by increased density of the crust below. These hollows may have been formed during the breaking up of the crust and subsequent removal of portions by currents, and not chiefly by the contraction of the crust. Thus the deflections at the stations (1) and (12) towards the land may be sufficiently accounted for, even if the land about Barcelona and Tornea does not rise sufficiently high to produce them. The deflection at station (2) is small. It seems probable that even if the North Sea has been produced according

to the theory of contraction of the crust, the parts near Dunkirk may have been somewhat hollowed out by the scouring of the tide through the Straits of Dover, so as to give the land, low as it is, every advantage in deflecting the plumb-line south. I have no means of knowing the character of the ground north of station (7) on the coast of Dorset. There is no difficulty, however, in accounting for the north deflection at that place, and even for a greater deflection, if the attraction of the country north of it is as much as the attraction of the land on Burleigh Moor on the north coast of Yorkshire. To this station I shall revert. With regard to stations (9) and (10), I gather the following information from the Ordnance Survey Volume. "At present there are no sufficient data for calculating exactly the disturbance" at Cowhythe (p. 662). It is supposed not to exceed 6" (p. 664); but the calculation is not made for any part of the mountains further south than 50 miles. The south deflection to be accounted for, viz.  $-3''.11$ , may in part be thus explained; or, even if, as before, the North Sea is supposed to have been formed by the contraction of the crust, the confined portion between the north coast of Aberdeen and the Orkney Islands may have been formed by the removal of the superficial strata by currents so as to produce a deficiency of attracting matter. So with respect to the other station, Ben Hutig, the unaccounted-for deflection, which is much smaller, viz.  $-0.60$ , may be easily explained, as the effect of the land has not been calculated further off than about 3 miles (pp. 660, 661). Thus, on the whole, the deflections at those coast-stations, where it is towards the land, can be pretty well accounted for, without calling in aid the deficiency of attraction of water and supposing that the crust below the ocean is not condensed.

The seven coast-stations of the second list, where the deflection is towards the sea, seem to bear individual testimony to the truth of the theory, that the crust below the ocean must have undergone greater contraction than other parts of the crust. The four stations (3), (4), (5), (6) on the south coast of the Isle of Wight all have deflections southwards; and their magnitudes diminish in the order that the distances from the sea increase,—that order being (3) High Port Cliff, (5) Boniface Down, (4) Week Down, (6) Dunnose (see the Contour Map of Isle of Wight in the volume of Plates accompanying the Ordnance Survey Volume). The amounts of the deflection seem almost to prove too much for the theory. Still they are all *in the direction* of the ocean, and seem certainly to indicate that there is a redundancy of matter, and not a deficiency, in that direction. Blackdown (7) is somewhat further inland than Dunnose is. If, then, the ocean and crust together do really produce the outstanding deflection southward at Dunnose, we shall have to suppose that the north deflection at Blackdown in the first list of coast-stations, arising from the land, is not much less than  $2.76 + 1.73 = 4.49$ , which is a little less than the calculated deflection at Burleigh Moor on the coast of Yorkshire, and is therefore not an unlikely amount. The other three coast-stations, (8), (11), (13), all bear out the

theory: though the three deflections are all small, they are towards the sea, the largest of them being at Fuglencæs, which is very near to the North Cape, and has a large expanse of ocean above it.

19. The least that can be gathered from the deflections of these coast-stations is, that they present no obstacle to the theory so remarkably suggested by the facts brought to light in India, viz. that mountain-regions and oceans on a large scale have been produced by the contraction of the materials, as the surface of the earth has passed from a fluid state to a condition of solidity—the amount of contraction beneath the mountain-region having been less than that beneath the ordinary surface, and still less than that beneath the ocean-bed, by which process the hollows have been produced into which the ocean has flowed. In fact the testimony of these coast-stations is in some degree directly in favour of the theory, as they seem to indicate, by *excess* of attraction towards the sea, that the contraction of the crust beneath the ocean has gone on increasing in some instances still further since the crust became too thick to be influenced by the principles of floatation, and that an additional flow of water into the increasing hollow has increased the amount of attraction upon stations on its shores.

Murree, Punjab,  
August 20, 1863.

June 2, 1864.

The Annual Meeting for the Election of Fellows was held this day.

Major-General SABINE, President, in the Chair.

The Statutes relating to the Election of Fellows having been read, General Boileau and Sir Andrew Scott Waugh were, with the consent of the Society, nominated Scrutators to assist the Secretaries in examining the lists.

The votes of the Fellows present having been collected, the following gentlemen were declared duly elected into the Society:—

Sir Henry Barkly, K.C.B.  
William Brinton, M.D.  
T. Spencer Cobbold, M.D.  
Alexander John Ellis, Esq.  
John Evans, Esq.  
William Henry Flower, Esq.  
Thomas Grubb, Esq.  
Sir John Charles Dalrymple Hay,  
Bart.

William Jenner, M.D.  
Sir Charles Locock, Bart., M.D.  
William Sanders, Esq.  
Col. William James Smythe, R.A.  
Lieut.-Col. Alexander Strange.  
Robert Warington, Esq.  
Nicholas Wood, Esq.

June 9, 1864.

Major-General SABINE, President, in the Chair.

Mr. W. Sanders; Mr. R. Warington; Dr. Jenner; Mr. J. Evans;